
“Introduction to Geostrophic Turbulence”

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Marcel Oliver

From 2D turbulence to MOLES

1. The closure problem
2. 2D single layer turbulence
3. Two-layer stratified turbulence
4. LES closures



1. The closure problem

Rotating Boussinesq equations

$$\begin{aligned}\partial_t \mathbf{u} + \nabla \cdot (\mathbf{u} \otimes \mathbf{u}) + 2\boldsymbol{\Omega} \times \mathbf{u} + \rho^{-1} \nabla p &= \mathbf{F} + D\mathbf{u} \\ \nabla \cdot \mathbf{u} &= 0\end{aligned}$$

Apply linear filter operation

$$\begin{aligned}\partial_t \bar{\mathbf{u}} + \bar{\nabla} \cdot (\bar{\mathbf{u}} \otimes \bar{\mathbf{u}}) + 2\boldsymbol{\Omega} \times \bar{\mathbf{u}} + \rho^{-1} \bar{\nabla} \bar{p} &= \mathbf{R}(\mathbf{u}) + \bar{\mathbf{F}} + \bar{D}\bar{\mathbf{u}} \\ \bar{\nabla} \cdot \bar{\mathbf{u}} &= 0\end{aligned}$$

with *eddy source term*

$$\mathbf{R}(\mathbf{u}) = \bar{\nabla} \cdot (\bar{\mathbf{u}} \otimes \bar{\mathbf{u}}) - \overline{\nabla \cdot (\mathbf{u} \otimes \mathbf{u})} + \overline{D\mathbf{u}} - \bar{D}\bar{\mathbf{u}}$$



2. Barotropic vorticity equation

$$\begin{aligned}\partial_t \zeta + \nabla^\perp \psi \cdot \nabla \zeta &= F + D_i \zeta + D_u \zeta \\ \zeta &= \Delta \psi\end{aligned}$$

Inviscid undamped equation conserves energy and enstrophy

$$E = \int |\nabla \psi|^2 dx \quad \text{and} \quad Z = \int |\zeta|^2 dx$$

Fourier representation

$$\partial_t \zeta_{\mathbf{k}} + \frac{1}{2\pi} \sum_{\mathbf{k}=\mathbf{p}+\mathbf{q}} \frac{\mathbf{p}^\perp \cdot \mathbf{q}}{p^2} \zeta_{\mathbf{p}} \zeta_{\mathbf{q}} = D_i(\mathbf{k}) \zeta_{\mathbf{k}} + D_u(\mathbf{k}) \zeta_{\mathbf{k}} + F_{\mathbf{k}}$$

Rate of energy transfer for mode $\mathbf{k}$:

$$\partial_t E_{\mathbf{k}} = \sum_{\{\mathbf{p}, \mathbf{q}\}: \mathbf{k}+\mathbf{p}+\mathbf{q}=0} T(\mathbf{k}|\mathbf{p}\mathbf{q}) + 2 D_i(\mathbf{k}) E_{\mathbf{k}} + 2 D_u(\mathbf{k}) E_{\mathbf{k}} + P_{\mathbf{k}}$$

where

$$T(\mathbf{k}|\mathbf{p}\mathbf{q}) = \frac{1}{2\pi} \mathbf{p}^\perp \cdot \mathbf{q} (q^2 - p^2) \operatorname{Re}[\psi_{\mathbf{k}} \psi_{\mathbf{p}} \psi_{\mathbf{q}}]$$



3. Resonant triad interactions (Fjørtoft, 1953)

Resonant triad: $\mathbf{k} + \mathbf{p} + \mathbf{q} = 0$

$$T(\mathbf{p}|\mathbf{k}\mathbf{q}) = -\frac{q^2 - k^2}{q^2 - p^2} T(\mathbf{k}|\mathbf{p}\mathbf{q})$$

and

$$T(\mathbf{q}|\mathbf{k}\mathbf{p}) = -\frac{k^2 - p^2}{q^2 - p^2} T(\mathbf{k}|\mathbf{p}\mathbf{q})$$

Energy transfer constraints

Suppose

$$p < k < q$$

If center mode $\mathbf{k}$ loses energy, modes $\mathbf{p}$ and $\mathbf{q}$ gain energy, and vice versa



4. Folklore

Energy goes upscale
Enstrophy goes downscale

Questions

- Can we prove it?
- Under which conditions?
- And in the real ocean/atmosphere?



4.1. “Proof” 1: Transient dynamics of spectral peak

Mean of spectral distribution (“Energy wavenumber”)

$$k_e = \frac{1}{E} \sum_k k E(k)$$

Variance of spectral distribution

$$I = \sum_k (k - k_e)^2 E(k) = Z - k_e^2 E$$

Conservation of energy and enstrophy implies:

$$\frac{dI}{dt} = -E \frac{dk_e^2}{dt}$$

As energy variance increases, the energy mean goes upscale



4.2. “Proof” 2: Infinite ranges

- Forcing scale k_f
- Infrared dissipation scale k_i
- Ultraviolet dissipation scale k_u

$$k_i \ll k_f \ll k_u$$

Forcing/dissipation balance

$$\begin{aligned}\varepsilon &= \varepsilon_i + \varepsilon_u \\ \eta &= k_f^2 \varepsilon = \eta_i + \eta_u\end{aligned}$$

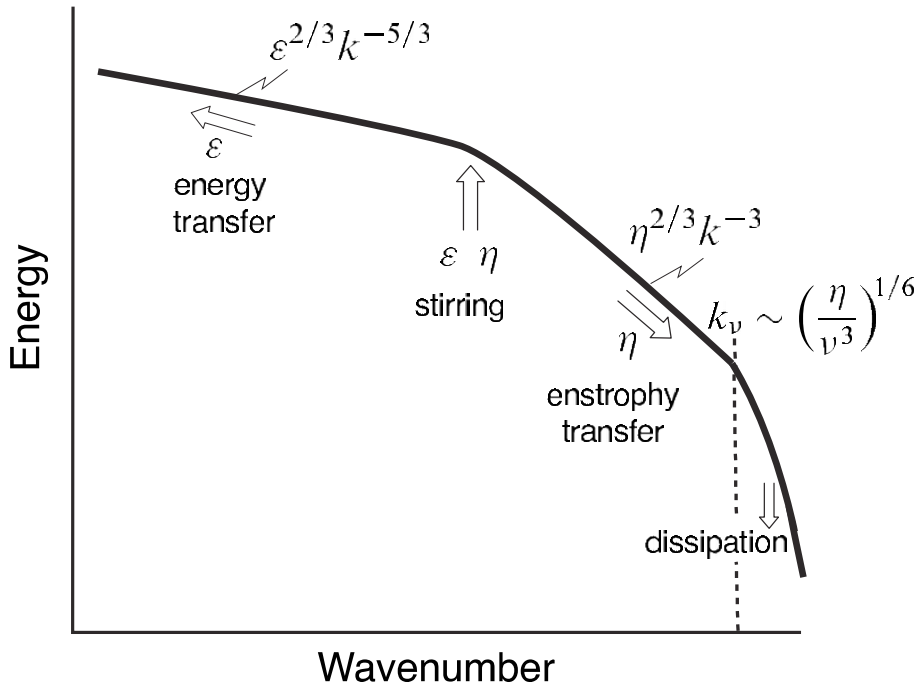
Since $\eta_i \leq k_i^2 \varepsilon_i$, we estimate

$$\eta \geq \eta_u = k_f^2 \varepsilon - \eta_i \geq k_f^2 \varepsilon - k_i^2 \varepsilon_i \geq (k_f^2 - k_i^2) \varepsilon = (1 - k_i^2/k_f^2) \eta$$

Thus, $\eta_u \rightarrow \eta$ in the limit $k_i/k_f \rightarrow 0$. Similarly, $\varepsilon_i \rightarrow \varepsilon$ when $k_f/k_u \rightarrow 0$.



5. Cascade picture for two-dimensional turbulence



6. Kolmogorov: Energy cascade

$$E_{[\kappa, 2\kappa]} = \int_{\kappa}^{2\kappa} S(k) \, dk \equiv g(\varepsilon, \kappa)$$

Rescaling time $t = Tt'$ and space $x = Lx'$, we demand

$$E'_{\kappa', 2\kappa'} = g(\varepsilon', \kappa') \equiv g(1, 1)$$

so that $\varepsilon = L^2 T^{-3} \varepsilon'$, $\kappa = L^{-1} \kappa'$, and therefore

$$E_{[\kappa, 2\kappa]} = \frac{L^2}{T^2} E'_{\kappa', 2\kappa'} = g(1, 1) \frac{\varepsilon^{2/3}}{\kappa^{5/3}}$$

Differentiating, we find

$$S(\kappa) - S(2\kappa) = \frac{2 g(1, 1)}{3} \frac{\varepsilon^{2/3}}{\kappa^{5/3}}$$

so that

$$S(\kappa) = S(2^{m+1}\kappa) + \frac{2 g(1, 1)}{3} \left(1 + \frac{1}{2^{5/3}} + \frac{1}{2^{10/3}} + \dots \right) \frac{\varepsilon^{2/3}}{\kappa^{5/3}} \implies$$

$$S(k) \sim c \frac{\varepsilon^{2/3}}{k^{5/3}}$$



7. Kraichnan–Batchelor–Leith: Enstrophy cascade

$$E_{[\kappa, 2\kappa]} = \int_{\kappa}^{2\kappa} S(k) \, dk \equiv g(\eta, \kappa)$$

Rescaling time $t = Tt'$ and space $x = Lx'$, we demand

$$E'_{\kappa', 2\kappa'} = g(\eta', \kappa') \equiv g(1, 1)$$

so that $\eta = T^{-3} \eta'$, $\kappa = L^{-1} \kappa'$, and therefore

$$E_{[\kappa, 2\kappa]} = \frac{L^2}{T^2} E'_{\kappa', 2\kappa'} = g(1, 1) \frac{\eta^{2/3}}{\kappa^2} \implies S(k) \sim c \frac{\eta^{2/3}}{k^3}$$

Dissipation length scale

$$\eta = \int_0^{k_\nu} (\nu k^2) (k^2) S(k) \, dk = c \nu \eta^{2/3} \int_0^{k_\nu} k \, dk = \tilde{c} \nu \eta^{2/3} k_\nu^2$$

so that

$$k_\nu = C \frac{\eta^{1/6}}{\nu^{1/2}}$$



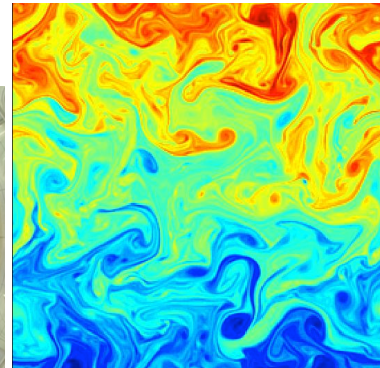
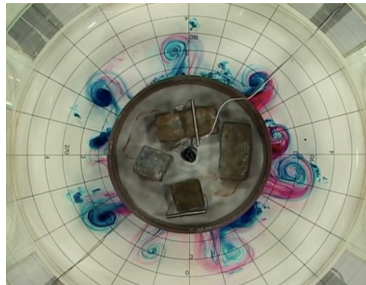
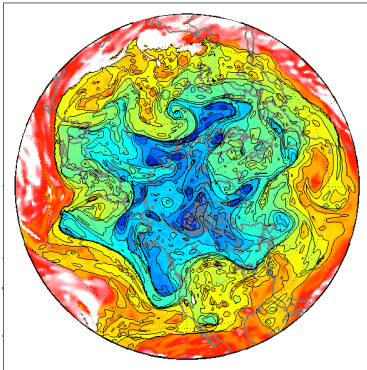
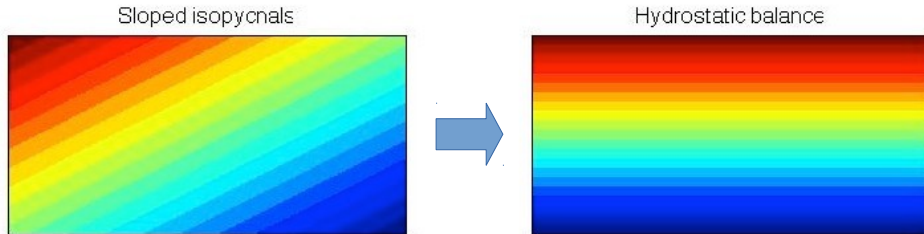
8. Limitations of the cascade picture

In reality

- Forcing is not spectrally localized (baroclinic instability!)
- Dissipation is not spectrally localized
- Forcing is too close to the grid scale
- Non-local triads can be significant



9. Baroclinic instability



Graphics from Sheane Keating,

<https://www.climate-science.org.au/sites/default/files/Baroclinic%20Instability%20Keating%20.pdf>



10. Two-layer QG as a model for ocean turbulence

Basic state: Uniform vertical shear

$$\psi_1^{\text{equil}} = -Uy \quad \text{and} \quad \psi_2^{\text{equil}} = 0$$

Eddy stream functions

$$\psi_1 = -yU + \psi_1^{\text{eddy}} \quad \text{and} \quad \psi_2 = \psi_2^{\text{eddy}}$$

Eddy barotropic and baroclinic stream functions

$$\psi = \frac{\psi_1^{\text{eddy}} + \psi_2^{\text{eddy}}}{2} \quad \text{and} \quad \tau = \frac{\psi_1^{\text{eddy}} - \psi_2^{\text{eddy}}}{2}$$

and associated PVs

$$q = \Delta\psi \quad \text{and} \quad \omega = \Delta\tau - k_d^2 \tau$$



10.0.1. Two-layer QG, ctd.

$$\begin{aligned}\partial_t q + [\psi, q] + [\tau, \omega] + \frac{U}{2} \partial_x (q + \Delta \tau) &= \frac{1}{2} D_i (\psi - \tau) + D_u \psi \\ \partial_t \omega + [\psi, \omega] + [\tau, q] + \frac{U}{2} \partial_x (\omega + q + k_d^2 \psi) &= -\frac{1}{2} D_i (\psi - \tau) + D_u \tau + \kappa \tau\end{aligned}$$

Energies (by mode)

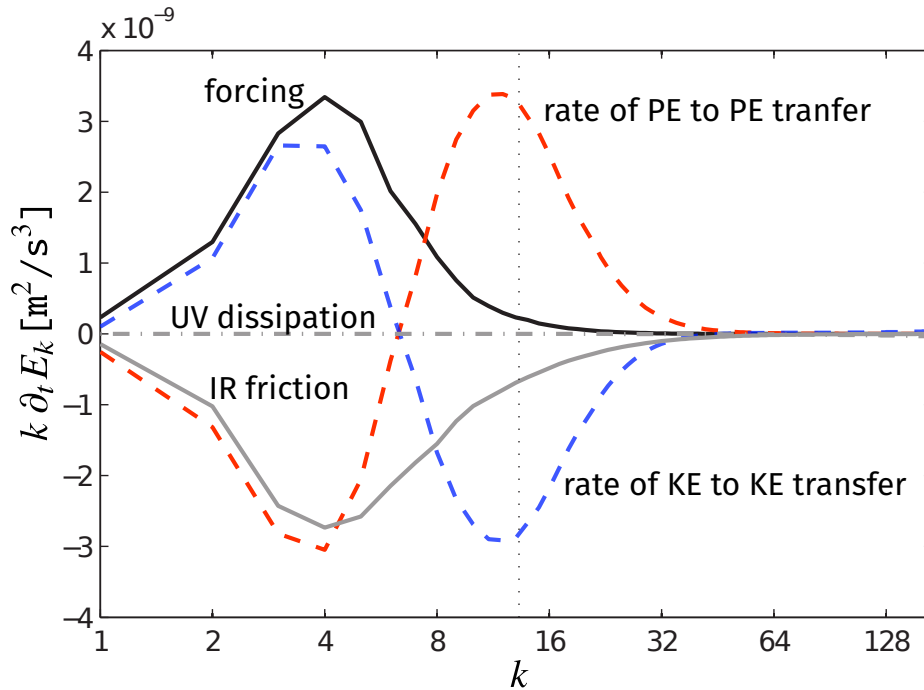
$$E^\psi = -\frac{1}{2} \sum_{\mathbf{k}} k^2 |\psi_{\mathbf{k}}|^2 \quad \text{and} \quad E^\tau = \frac{1}{2} \sum_{\mathbf{k}} (k^2 + k_d^2) |\tau_{\mathbf{k}}|^2$$

Energies (kinetic and potential)

$$E^K = -\frac{1}{2} \sum_{\mathbf{k}} k^2 (|\psi_{\mathbf{k}}|^2 + |\tau_{\mathbf{k}}|^2) \quad \text{and} \quad E^P = \frac{1}{2} \sum_{\mathbf{k}} k_d^2 |\tau_{\mathbf{k}}|^2$$



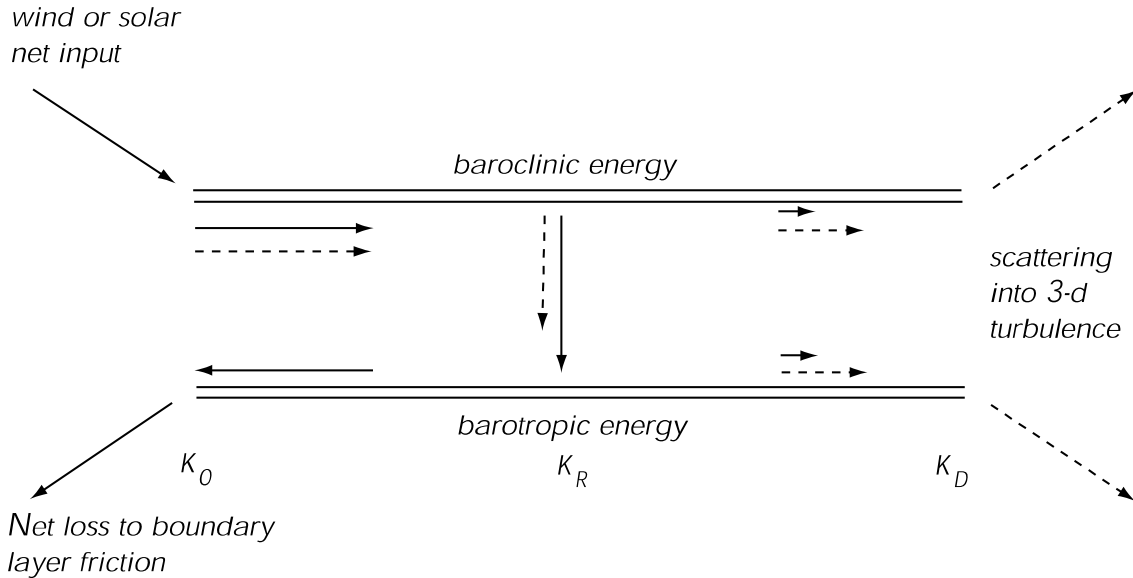
10.0.2. Two-layer QG, numerical energy transfer rates



Graphics adapted from Jansen/Held (2014)



10.1. Energy transfers in two-layer geostrophic turbulence



Graphics from Vallis (2005)



10.2. How does this apply to the ocean?

- Upper ocean: $N \approx 0.0025\text{s}^{-1}$, $H \approx 4\text{ km}$, $f_0 \approx 10^{-4}\text{s}^{-1}$
- Baroclinic Rossby radius: $L_d \approx 100\text{ km}$
- Most unstable scale: approx. $4 L_d \approx 400\text{ km}$
- Locally much smaller values possible
- At wave numbers larger than k_d , ageostrophic effects come in
- At wave numbers around k_d , eddy viscosity comes in
- There are also surface modes, possibly higher baroclinic modes



11. Viscous parameterization (Leith)

Idea 1: Determine “eddy viscosity” ν_* such that a given dissipation wavenumber, *within the inertial range*, maintains the rate of enstrophy dissipation:

$$k_* = C \frac{\eta^{1/6}}{\nu_*^{1/2}}$$

On the other hand, from barotropic vorticity equation,

$$\begin{aligned} \frac{1}{2} \frac{d}{dt} \zeta^2 + \frac{1}{2} \nabla \cdot (\mathbf{u} \zeta^2) &= \nu_* \zeta \Delta \zeta && + \text{IR friction} + \text{forcing} \\ &= -\nu_* |\nabla \zeta|^2 + \nu_* \nabla \cdot (\zeta \nabla \zeta) && + \text{IR friction} + \text{forcing} \end{aligned}$$

Thus, area-mean-rate of enstrophy dissipation $\eta = \nu_* \langle |\nabla \zeta|^2 \rangle$

Idea 2: Swap local and global values

$$\nu_* = \left(\frac{C}{k_*} \right)^3 \langle |\nabla \zeta|^2 \rangle^{\frac{1}{2}} = \left(\frac{C}{k_*} \right)^3 |\nabla \zeta|$$

→ MOLES (Fox–Kemper and Menemenlis, 2008)



12. Outlook

- Discrete, small-stencil operators (WIP)
- Backscatter (Stephan + Anton's talks later)
- Can we get off the cascade picture? (Non-local triads!)
- Stochastic closures (M2 and others)
- Wave-turbulence interaction (IDEMIX!)

